

U SNO R obotic A strometric T elescope

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Operations lead: Charlie Finch

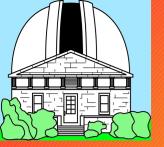
Astrometry Department

U.S. Naval Observatory



layout of talk

- History: plan and construction
- URAT operations north (NOFS)
- URAT operations south (CTIO)
- Results, Catalogs
- Future: what is next?



History: project plan and construction of the URAT instrument





USNO Instrument Shop(s)

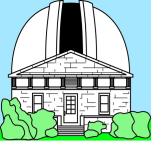
- in Washington, DC
 - lead by Gary Wieder, latest machines
 - design and build astronomical instrumentation
 - highly regarded (NPOI, SDSS, URAT, clocks)
- at NOFS (Naval Observatory Flagstaff Station)
 - lead by Mike DiVittorio
 - maintenance of telescopes and instrumentation
 - adaptive optics design, NPOI





URAT project

- Goal #1: follow-up of UCAC in pre-Gaia era
 - go deeper, many more stars
 - more precise: many sky overlaps
 - more accurate: new detectors, no CTE problem
 - get parallaxes of nearby stars without pre-selection
- Solution: "red lens" with modified astrograph
 - 4-shooter camera: each 10,560 x 10,560 pixel
 - each exposure covers 28 sq.deg (about 1 GB)
 - use UCAC reference stars, survey to R = 18 mag



URAT project (cont.)

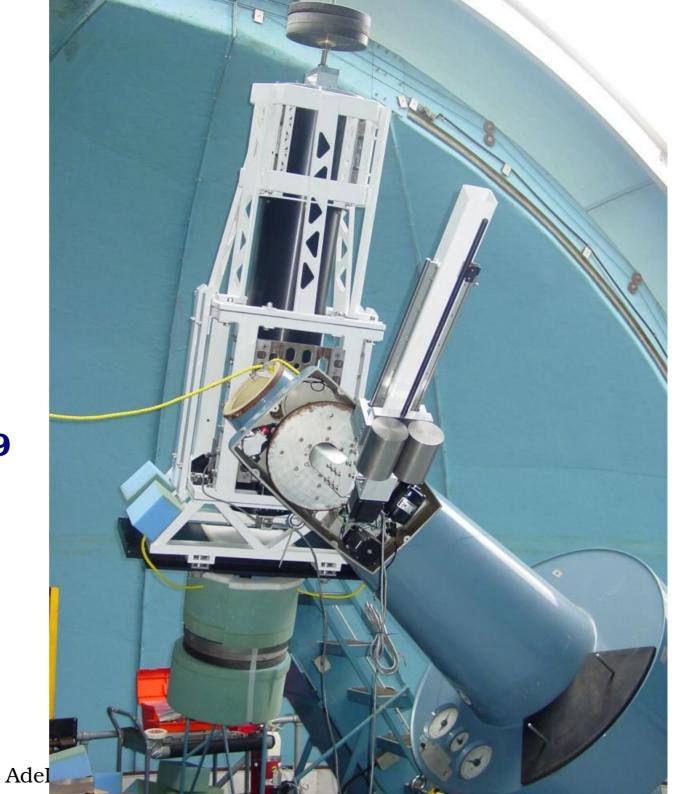
- Goal #2: bright stars = supplement Gaia
 - use new neutral density spot + objective grating
 - access all bright stars up to Sirius

Milestones

- 10k test camera: 1st light October 2007
- funding for 4-shooter camera in FY2008
- 1st light in Washington DC: Sept. 2011
- operations at NOFS: April 2012 June 2015
- begin operations at CTIO: Oct. 2015

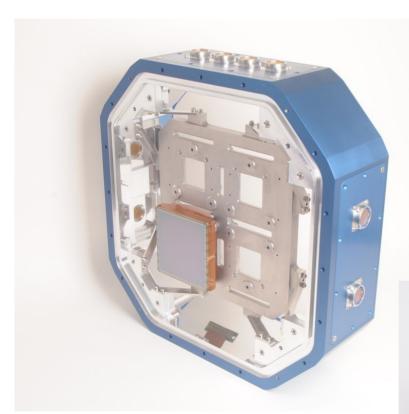


astrograph
May 2009
at USNO

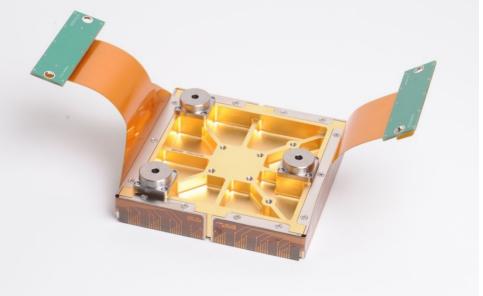


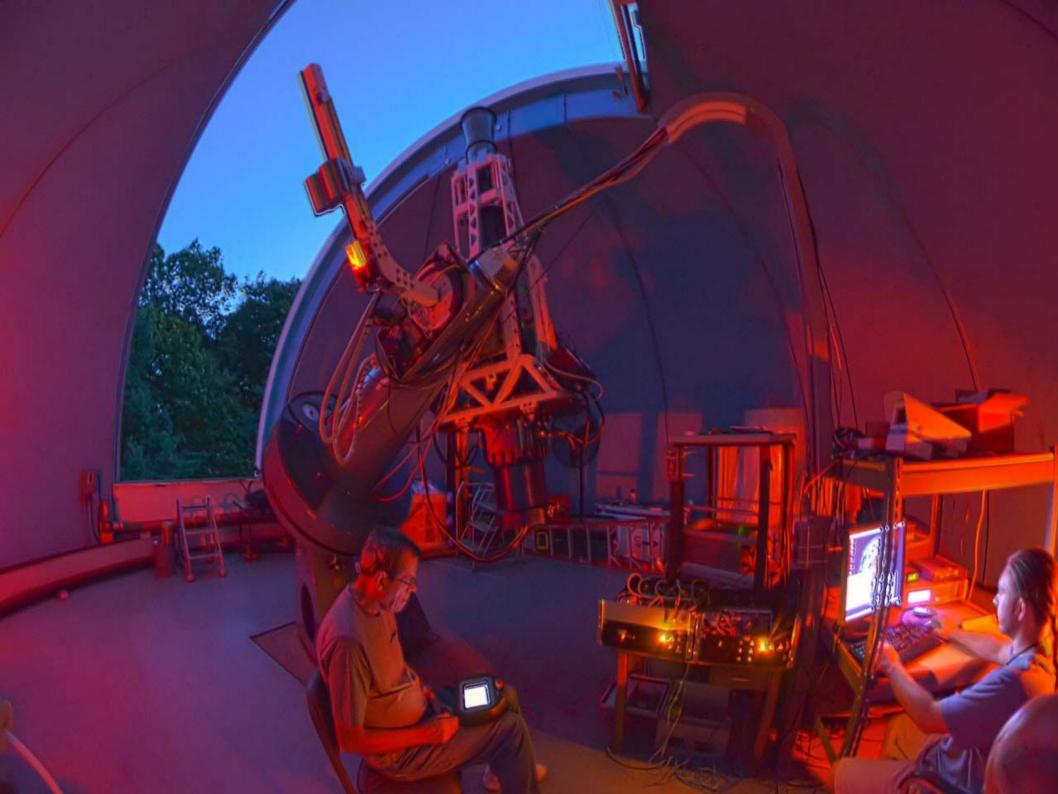


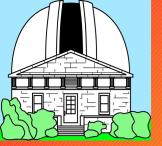
spring 2010, 10k packaging



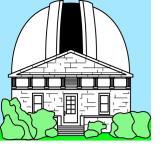








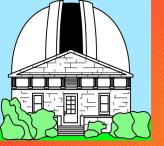
URAT operations: Northern Hemisphere NOFS, Arizona



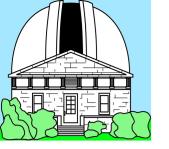
URAT observations north

- complete re-make of astrograph 2008-2010
- Naval Obs. Flagstaff Station: 2012-2015
- 28 sq. deg. per exposure! (4 detectors)
- 10 mas per image (well exposed stars)
- multiple sky overlaps / year, 7 18 mag
 - clocked anti-blooming: extend dynamic range
 - around full Moon: short expos. grating: 3 15 mag
- solve for positions, motions + parallax





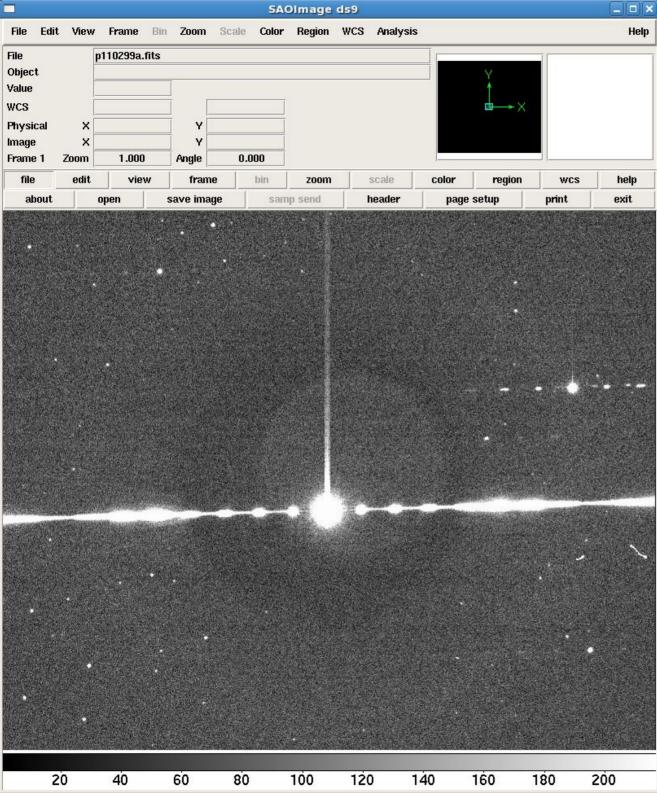
URAT operations: Southern Hemisphere CTIO, Chile



URAT operations south

- by 2014 known: Gaia will be successful:
 - shift from goal #1 to #2: bright stars only
- move instrument
 - pack up at NOFS (June 2015)
 - ship dewar window out to get new ND spot
 - deploy at CTIO (Sept / Oct 2015)
- 10, 30, 60 sec exposures:
 - general, shallow, quick survey: 3.5 to 17 mag
 - targeted observations: -1.5 to 4.5 mag stars



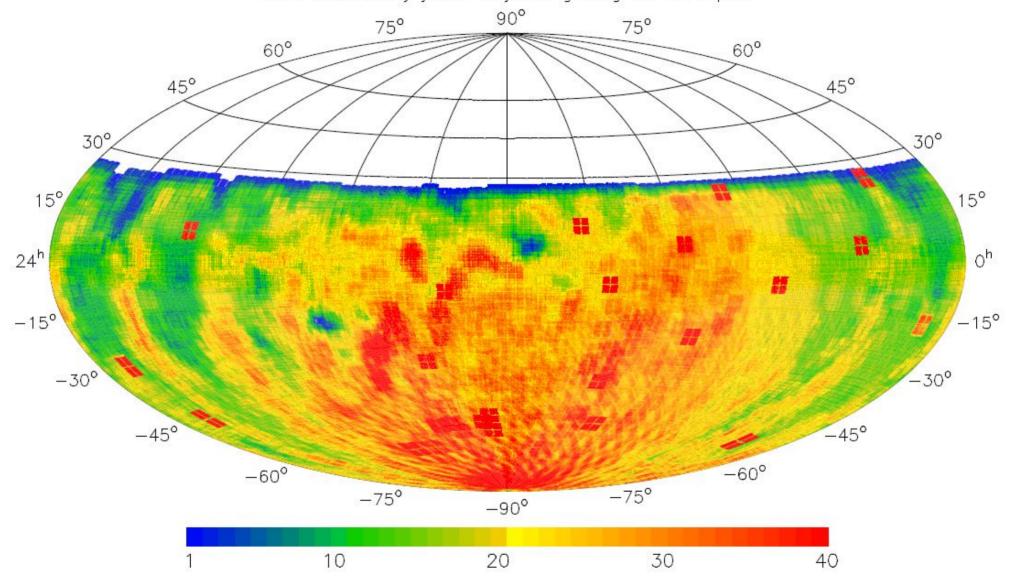


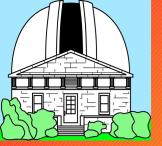
5 sec
exposure of
Sirius with
ND spot filter
and grating



observing progress mid Aug 2016







Results: published catalogs data mining efforts



products

- URAT1 = public star catalog (north)
 - based on first 2 years of operations at NOFS
 - typically 5 to 30 mas precision = f (mag)
 - preliminary proper motions with 2MASS
- UPC = URAT parallax catalog (Finch+2016)
 - over 100,000 entries = largest since Hipparcos
 - precision on 3 to 12 mas level = f(obs.history)
- URAT2 = all data from 3 years at NOFS
 - not public, internal use only at this point

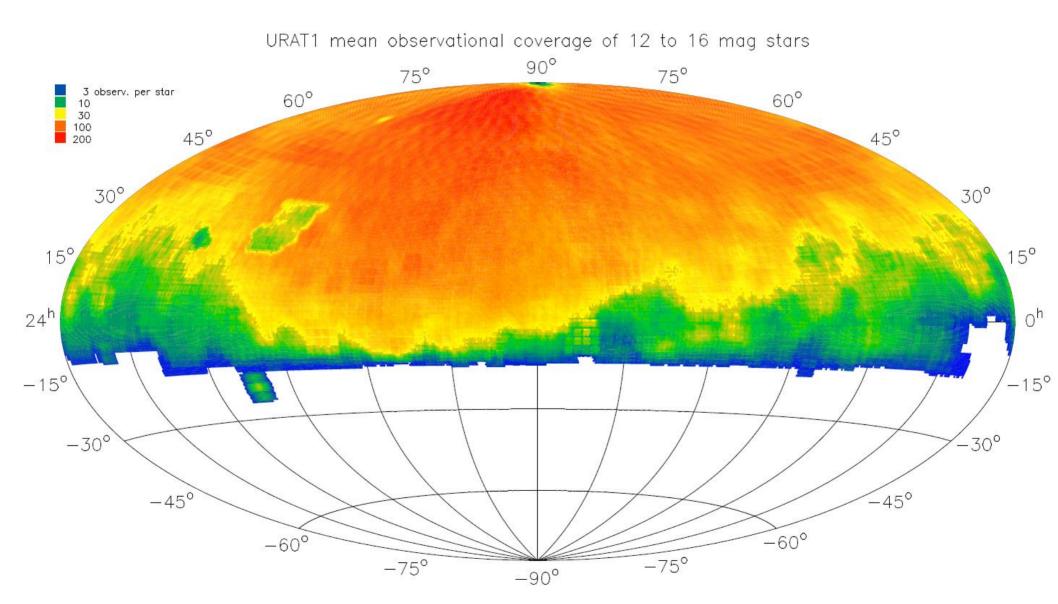


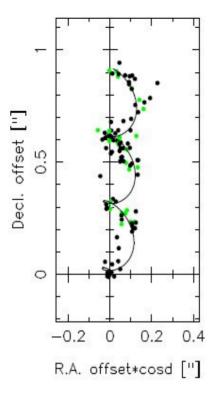
URAT1 catalog (2015)

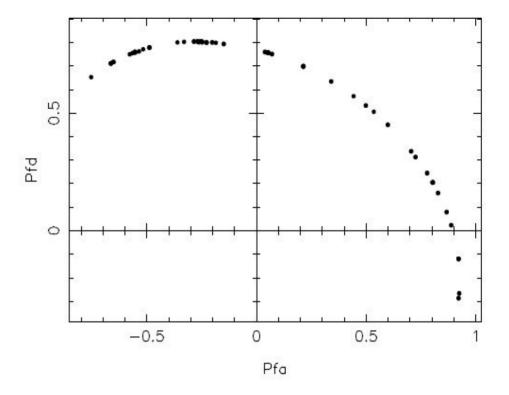
```
total numb. URAT1 stars = 228276482
number stars with 1 obs = 10309229
number stars with 2 obs = 8875122
average numb. obs/star = 24.3
                                   82.64%
number valid 2MASS data = 188656145
no 2MASS match stars = 39620337
                                    of
stars >= 3 obs., no 2MASS= 39079551 URAT1
number valid APASS data = 37010348
                                   16.21%
APASS stars valid B mag = 29313850
                                   of
APASS stars valid V mag = 30057593 URAT1
APASS stars valid g mag = 32340624 stars
APASS stars valid r mag = 32474206
APASS stars valid i mag = 28052917
```

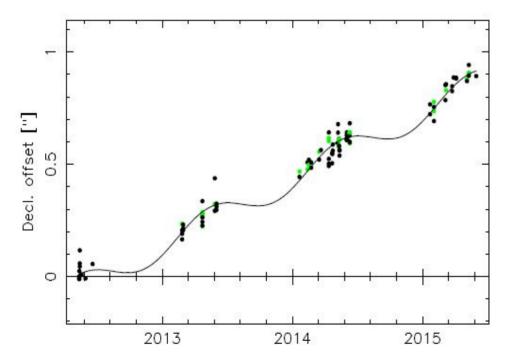


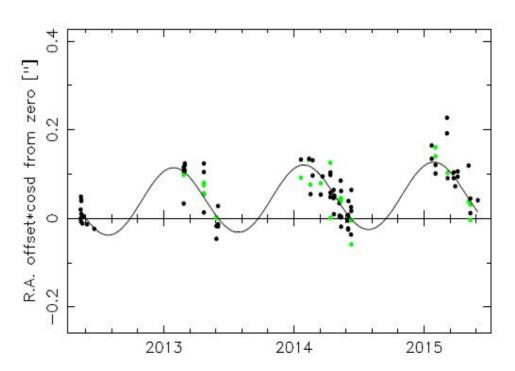
URAT1 sky coverage

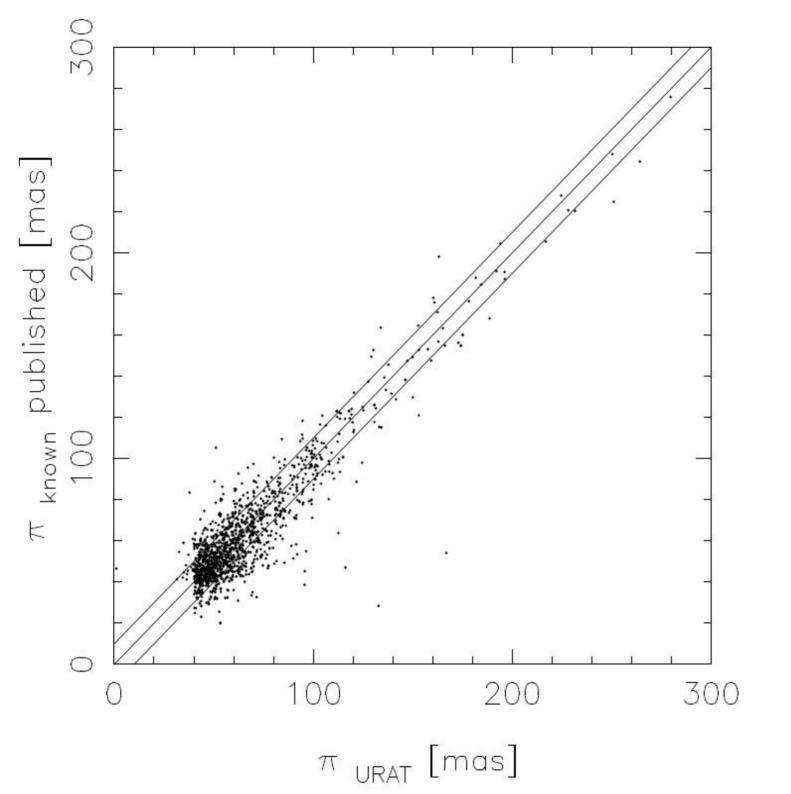




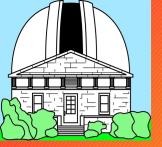








UPC = **URAT Parallax** Catalog (2016)112,177 stars



Future: what is next? (life after Gaia DR1)



anything useful?

- Accurate (5-10 mas) positions of bright stars not yet included in Gaia DR (<= 4 mag)
- Continue URAT observing at CTIO until about mid 2017
- Data reduction with Gaia reference stars is in progress
- URAT is useful for time-domain astronomy: possible collaboration with ZTF? Others?



summary

- URAT = follow-up to successful UCAC: bright + faint
- URAT1: 3-18 mag, 5-30 mas, NOFS, 228 mill. stars
- UPC = URAT Parallax Catalog (112,000 stars)
- Next:
 - data mining of full 3-year northern observ.: AGNs ...
 - very bright stars not published by Gaia yet
 - phase out observing at CTIO by mid 2017
- New project: time-domain astronomy (4 to 17 mag) ??